

Frequency-domain self-force scatter calculations using the Gegenbauer procedure

Chris Whittall
University of Birmingham

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UNIVERSITY OF
BIRMINGHAM

GRAVITATIONAL
WAVE ASTRONOMY

Gravitational scattering

Why study scattering?

- Clean, well-defined asymptotic in/out states.
- Probe strong-field (sub-ISCO) at low energy.
- EOB Hamiltonian H_{EOB} completely determined by scatter angle $\chi(E, L)$. [Damour 2016+]
- B2B maps between scatter and bound orbit. [Kalin & Porto 2020+]

Complementary approaches include:

- **Numerical relativity:** strong-field, low-separation, even mass ratios.
[See talks by S. Swain, O. Long tomorrow.]
- **Post-Minkowskian expansion (expansion in G):** weak-field, arbitrary mass-ratio, analytical.
- **Gravitational self-force (expansion in mass-ratio ϵ):** strong- and weak-field, small mass-ratio, numerical.

Interest in developing numerical GSF calculations along scatter orbits.

Scalar-field self-force and scattering

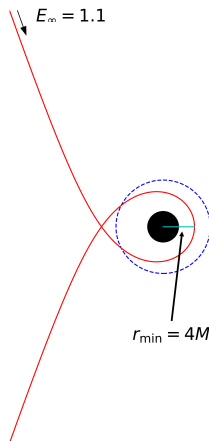
- Scalar charge q with mass μ scattered off a Schwarzschild black hole of mass M . Scalar field:

$$\nabla\Phi = -4\pi q \int \frac{\delta^4(x^\alpha - x_p^\alpha(\tau))}{\sqrt{-g}} d\tau,$$

where $\epsilon := q^2/\mu M \ll 1$ is the expansion parameter.

- At leading order take $x_p^\alpha(\tau)$ to be a scatter geodesic: parameterised by energy at infinity, E_∞ , and periapsis radius r_{\min} .
- Particle feels a self-force due to interaction with its own scalar field:

$$u^\beta \nabla_\beta (\mu u^\alpha) = q \nabla_\beta \Phi^R := F^\alpha.$$



Numerical self-force calculations

Decompose the field into spherical harmonics centred on the Schwarzschild black hole:

$$\Phi(t, r, \theta, \phi) = \frac{1}{r} \sum_{\ell=0}^{\infty} \sum_{m=-\ell}^{+\ell} \psi_{\ell m}(t, r) Y_{\ell m}(\theta, \phi)$$

Self-force can be computed using mode-sum regularisation:

$$F_{\alpha}(\tau) = \sum_{\ell=0}^{\infty} \left\{ \nabla_{\alpha} \left[\frac{1}{r} \sum_{m=-\ell}^{+\ell} \psi_{\ell m} Y_{\ell m} \right]_{x_p^{\pm}(\tau)} \mp \left(\ell + \frac{1}{2} \right) \underbrace{A_{\alpha}(\tau) - B_{\alpha}(\tau)}_{\text{regularisation parameters}} \right\}$$

Time-domain

- Solve $(1+1)d$ **PDEs** for $\psi_{\ell m}(t, r)$.
- Scattering: [Long & Barack 2209.03740]

Frequency-domain

- Construct $\psi_{\ell m}(t, r)$ from frequency-modes, obtained by solving **ODEs**.
- Scattering: [Whittall & Barack 2305.09724]

Direct reconstruction

- Frequency-modes $\psi_{\ell m \omega}$ obey ODE

$$\frac{d^2 \psi_{\ell m \omega}}{dr_*^2} - [V_\ell(r) - \omega^2] \psi_{\ell m \omega} = S_{\ell m \omega}(r)$$

- Retarded solution expressed in terms of homogeneous solution basis $\psi_{\ell \omega}^\pm(r)$ using variation of parameters:

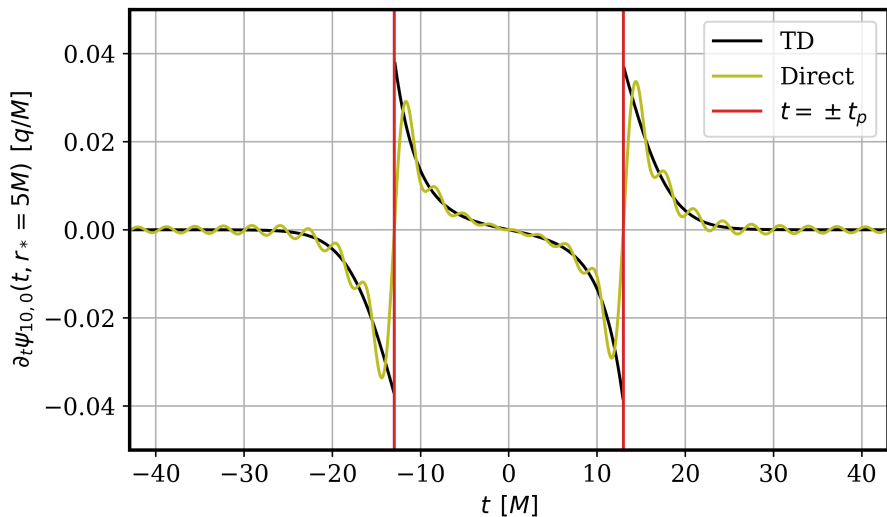
$$\psi_{\ell m \omega}(r) = \psi_{\ell \omega}^+(r) \int_{r_{\min}}^r \frac{\psi_{\ell \omega}^-(r') S_{\ell m \omega}(r')}{W_{\ell \omega} f(r')} dr' + \psi_{\ell \omega}^-(r) \int_r^{+\infty} \frac{\psi_{\ell \omega}^+(r') S_{\ell m \omega}(r')}{W_{\ell \omega} f(r')} dr'$$

- Direct reconstruction:

$$\psi_{\ell m}(t, r) \approx \Psi_{\ell m}(t, r; \omega_{\max}) := \int_{-\omega_{\max}}^{+\omega_{\max}} \psi_{\ell m \omega}(r) e^{-i\omega t}$$

is **not practical** due to the **Gibbs phenomenon**.

Direct reconstruction



EHS reconstruction [Barack, Ori & Sago 2008]

- **EHS reconstruction**: recover $\psi_{\ell m}(t, r)$ separately in $r \leq r_p(t)$ and $r \geq r_p(t)$ using homogeneous solutions.
- For example, field modes in the “internal” region $r \leq r_p(t)$ reconstructed from

$$\tilde{\psi}_{\ell m \omega}^{-}(r) := \psi_{\ell \omega}^{-}(r) \int_{r_{\min}}^{+\infty} \frac{\psi_{\ell \omega}^{+}(r') S_{\ell m \omega}(r')}{W_{\ell \omega} f(r')} dr'.$$

- Restores **exponential, uniform convergence**.
- **External reconstruction**: EHS cannot be applied in $r > r_p(t)$ for scatter orbits.
- **Cancellation problem**: significant loss of precision at high eccentricities [van de Meent 2016]. Limited to $r_p \sim r_{\min}$ for scatter orbits.

Gibbs-complementary reprojection

- Say that a basis of functions $\{C_k\}$ (on a compact time interval I) form a **Gibbs-complement** to the Fourier basis if: [Gottlieb & Shu 1998]
 - ① for any function f which is analytic on I , the expansion in the $\{C_k\}$ basis converges exponentially to f , **and**
 - ② the projection of the high-frequency Fourier content onto the low-degree C_k can be made exponentially small.
- **Gegenbauer polynomials** $\{C_k^\lambda(s)\}$ orthogonal on $[-1, 1]$ wrt weighted inner product,

$$\int_{-1}^{+1} (1 - s^2)^{\lambda-1/2} C_n^\lambda(s) C_m^\lambda(s) ds = h_n^\lambda \delta_{nm}.$$

- ▶ Generalisation of Legendre ($\lambda = 1/2$) and Chebyshev ($\lambda = 0, 1$) polynomials.
- ▶ **Gibbs-complementary** to the Fourier basis.

Gegenbauer reconstruction [Gottlieb & Shu 1992 (et al), 1994, 1995, 1997]

Suppose $\psi_{\ell m}(t, r)$ is analytic (at fixed r) on interval $a \leq t \leq b$:

- 1 Compute the partial Fourier integrals using the inhomogeneous modes $\psi_{\ell m \omega}(r)$ for $t \in [a, b]$:

$$\Psi_{\ell m}(t, r; \omega_{\max}) := \int_{-\omega_{\max}}^{+\omega_{\max}} \psi_{\ell m \omega}(r) e^{-i\omega t} d\omega.$$

- 2 Project onto the Gegenbauer basis:

$$g_k^\lambda(r; \omega_{\max}) := \frac{1}{h_k^\lambda} \int_{-1}^1 (1-s^2)^{\lambda-1/2} \Psi_{\ell m}(t(s), r; \omega_{\max}) C_k^\lambda(s) ds.$$

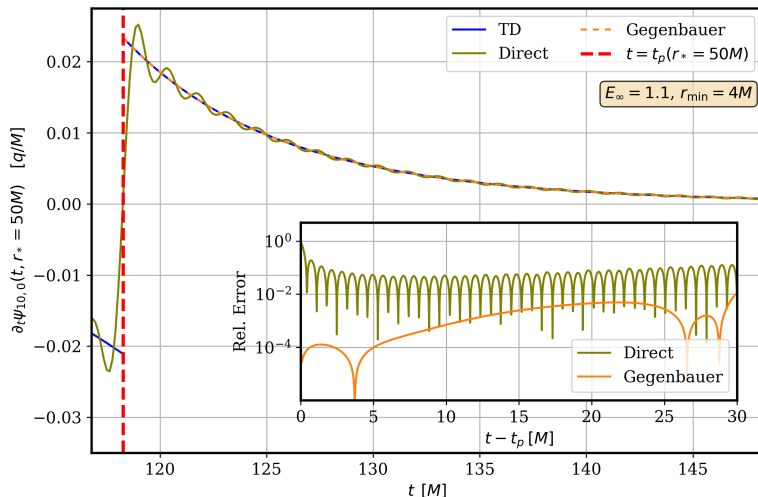
where $t(s) = [(b-a)s + (a+b)]/2$.

- 3 Approximate

$$\psi_{\ell m}(t, r) \approx \sum_{k=0}^N g_k^\lambda C_k^\lambda(s(t)).$$

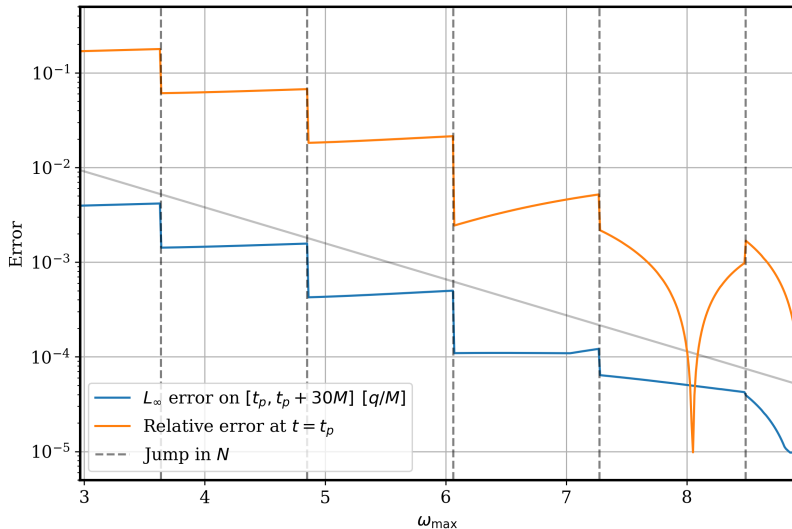
Gegenbauer approximant converges **uniformly** and **exponentially** on $a \leq t \leq b$, provided N, λ and $\omega_{\max} \rightarrow \infty$ in **linear proportion**.

Internal reconstruction: $r \leq r_p(t)$



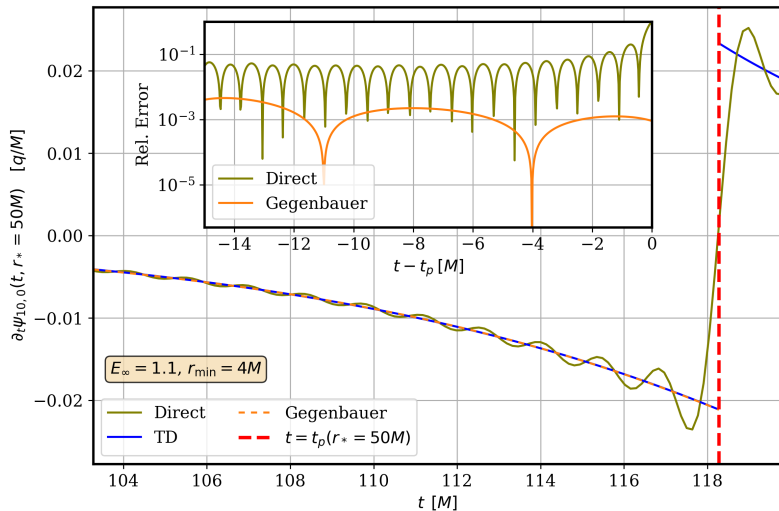
Gegenbauer reconstruction effective and outperforms Direct reconstruction.

Convergence



Convergence expected to be uniform and exponential.

External reconstruction: $r \geq r_p(t)$



Gegenbauer reconstruction enables calculations in $r > r_p(t)$.

Concluding remarks

- **Coming soon to arXiv:** proof of concept calculations for scalar-field modes $\psi_{\ell m}(t, r)$ with scattering source.
- Working towards full scalar-field self-force scatter calculation:
 - ▶ **Computational efficiency:** reconstruction, radial integrals in FD.
 - ▶ **Optimal choices of N/ω_{\max} and λ/ω_{\max} .**
- **Gravity** is the ultimate aim .
- Potential for improved methods and new applications.